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Dissipative inertial transport patterns near coherent Lagrangian eddies in the ocean

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Recent developments in dynamical systems theory have revealed long-lived and coherent Lagrangian (i.e., material) eddies in incompressible, satellite-derived surface ocean velocity fields. Paradoxically, observed drifting buoys and floating matter tend to create dissipative-looking patterns near oceanic eddies, which appear to be inconsistent with the conservative fluid particle patterns created by coherent Lagrangian eddies. Here, we show that inclusion of inertial effects (i.e., those produced by the buoyancy and size finiteness of an object) in a rotating two-dimensional incompressible flow context resolves this paradox. Specifically, we obtain that anticyclonic coherent Lagrangian eddies attract (repel) negatively (positively) buoyant finite-size particles, while cyclonic coherent Lagrangian eddies attract (repel) positively (negatively) buoyant finite-size particles. We show how these results explain dissipative-looking satellite-tracked surface drifter and subsurface float trajectories, as well as satellite-derived Sargassum distributions. © 2015 AIP Publishing LLC. [http://dx.doi.org/10.1063/1.4928693]

Satellite-tracked drifting buoy trajectories and satellite-derived algal distributions are commonly used in oceanography to infer Lagrangian aspects of the surface ocean circulation. At the same time, dynamical systems techniques applied to surface ocean velocities inferred from satellite altimetry reveal persistent coherent Lagrangian eddies. Paradoxically, buoys and algae display dissipative-looking patterns in contrast to the conservative-looking coherent Lagrangian eddies. Here, we show that the dissipative patterns are due to inertial effects superimposed on the conservative fluid patterns produced by coherent Lagrangian eddies.

I. INTRODUCTION

The work reported in this paper provides an explanation for the dissipative behavior of drifting buoys and floating matter on the ocean surface near coherent Lagrangian (i.e., material) eddies. Such eddies impose conservative behavior on nearby fluid particles in incompressible two-dimensional flows, which seems at odds with the observed dissipative patterns.

A revealing example of observed dissipative behavior is that of two RAFOs floats (acoustically tracked, subsurface drifting buoys of quasi isobaric type) in the southeastern North Pacific (Fig. 1). Initially close together, the two floats (indicated in red and green in Fig. 1) take significantly divergent trajectories on roughly the same depth level (320 m) relative to the floats positional uncertainty, which does not exceed 10 km (Garfield et al., 1999 and Collins et al., 2013). This behavior at first sight might be attributed to sensitive dependence of fluid particle trajectories on initial particle positions in a turbulent ocean. But, analysis of satellite altimetry measurements reveals that the floats on the date of closest proximity fall within a region of roughly 100-km radius characterized by a bulge of the sea surface height (SSH) field (selected isolines are indicated by dashed curves in Fig. 1).

FIG. 1. Trajectories of two RAFOs floats (red and green curves) and selected snapshots of a westward propagating bulge of the satellite altimetric sea surface height (SSH) field (dashed curves indicate selected isolines) in the southeastern North Pacific. The dots indicate the positions of the floats on the dates that the SSH bulge is shown.
This SSH bulge propagates westward at a speed slower than the geostrophically inferred clockwise tangential speed at its periphery, suggesting the presence of a mesoscale anticyclonic eddy capable of holding fluid (Chelton et al., 2011). Indeed, this SSH eddy may be identified with the surface manifestation of a California Undercurrent eddy; such eddies, referred to as “cuddies,” have been argued to be important transport agents (Garfield et al., 1999). However, while one float is seen to loop anticyclonically accompanying the eddy very closely, the other float anticyclonically spirals away from the eddy rather quickly, representing a puzzle.

Even more puzzling is that the two floats actually initially lie, as we show below, within the same coherent Lagrangian eddy (Beron-Vera et al., 2013 and Haller and Beron-Vera, 2013; 2014). Representing an elliptic Lagrangian coherent structure (LCS; cf. Haller, 2015), the boundary of such an eddy defies the exponential stretching of typical material loops in turbulence. In effect, the eddy in question exhibits minimal filamentation and deformation over several months and thus is expected to trap and carry within both floats.

The behavior of one of the floats supports this scenario and thus the altimetry-derived (i.e., geostrophic) upper-ocean current representation that sustains the coherent Lagrangian eddy. Ageostrophic processes of various types may be acting in the upper ocean, but these cannot explain the substantively different behavior of the other float. Indeed, ageostrophic effects cannot be so different on two initially nearby fluid particles in a region of mostly regular flow. Therefore, to resolve the puzzle, effects of a different class must be accounted for.

Here, we argue theoretically, and show both numerically and observationally, that such effects can be produced by inertia, i.e., buoyancy and size finiteness. Inertial effects are commonly considered in atmospheric transport studies. These range from studies aimed at explaining observed motion of meteorological balloons (Paparella et al., 1997; Provenzale, 1999; and Dvorin et al., 2001) and spread of volcanic ash (Haszpra and Tél, 2011), to theoretical and numerical studies of particulate matter dispersal (Haller and Sapsis, 2008; Sapsis and Haller, 2009; and Tang et al., 2009). In oceanography, inertial effects have been also taken into account in several problems including sedimentation (Nielsen, 1994), plankton sinking (Stommel, 1949), patchiness (Reigada et al., 2003), and selfpropulsion (Peng and Dabiri, 2009). However, they have been rarely considered in the motion of drifting buoys, macroscopic algae, or debris. To the best of our knowledge, their potential importance in influencing the motion of floats was noted only by Tanga and Provenzale (1994).

Our theoretical results reveal that while the boundary of a coherent Lagrangian eddy represents a transport barrier for fluid particles, it does not do so for inertial particles. Instead, a coherent Lagrangian eddy attracts or repels initially close inertial particles, depending on the particles’ density ratio with the ambient fluid and on the polarity (rotation sense) of the eddy.

We first validate these numerically using altimetry-derived currents in several regions of the ocean. Next, we use our findings to explain observed behavior in various ocean areas starting with the aforementioned floats, then proceeding with satellite-tracked surface drifting buoys (drifters), and finally macroscopic algae (Sargassum) distributions.

We emphasize that because our approach uses observation-based velocity, it enables feature matching and analysis of specific measurements. Furthermore, our approach is self-consistent within the realm of incompressible two-dimensional flows. This is in marked contrast with a previous approach to surface ocean pattern formation (Zhong et al., 2012), which considered passive advection by the surface velocity from a primitive-equation model (i.e., a truncation of the three-dimensional velocity). This is destined to create dissipative-looking patterns, but no actual passive tracer follows such a virtual velocity field.

We also note that our results are not applicable to the problem of accumulation of debris in subtropical gyres, which has been recently investigated by Froyland et al. (2014) using probabilistic methods. The so-called great ocean garbage patches are produced by convergent wind-induced Ekman transport (Maximenko et al., 2011). The Ekman dynamics governing basin-scale motions are very different than the quasigeostrophic dynamics governing mesoscale motions, our focus here. The former can produce dissipative patterns on the surface ocean by themselves, but the latter cannot unless inertial effects are taken into account, as we noted above and demonstrate below.

The remainder of the paper is organized as follows. Section II A presents the mathematical setup required to formally introduce the coherent Lagrangian eddy notion, which is briefly reviewed in Section II B. The theoretical results relating to behavior of inertial particles near coherent Lagrangian eddies are presented in Section II C. In Section II D, further insight into inertial particle motion is provided. Numerical validation of the theoretical results is presented in Section III. In Section IV, the theoretical results are used to explain observed behavior in the ocean. A summary and discussion is offered in Section V. Finally, Appendix A includes details of the asymptotic analysis leading to our theoretical results, Appendix B is reserved for the description of the several data sets employed, Appendix C gives some details of the various numerical computations performed, and Appendix D presents an extended analysis for the float data set.

II. THEORY

A. Mathematical setup

We consider an incompressible two-dimensional velocity field, \( v(x, t) \), where position \( x \) ranges on some open domain of \( \mathbb{R}^2 \) and time \( t \) is defined on a finite interval. Specifically, we consider

\[
\eta = \frac{g}{f} \nabla^+ \eta, \tag{1}
\]

where \( \eta(x, t) \) is the SSH; the constants \( f \) and \( g \) stand for Coriolis parameter (twice the local vertical component of the
Earth’s angular velocity) and acceleration of gravity, respectively; and \( \perp \) represents a 90°-anticlockwise rotation. The velocity field (1) is representative of quasigeostrophic motions in the upper ocean, i.e., characterized by a small Rossby number, \( \text{Ro} := V/L/f \), where \( L \) and \( V \) are typical length and velocity scales, respectively. In particular, (1) is suitable to investigate transport near mesoscale eddies, our focus here. Fluid particles evolve according to

\[
\dot{x} = v.
\]  

(2)

Let \( F^t_t(x_0) := x(t;x_0, t_0) \) be the flow map that takes time \( t_0 \) positions to time \( t \) positions of fluid particles obeying (2). An objective (i.e., frame-invariant) measure of material deformation in (2) is the right Cauchy–Green strain tensor

\[
C := (DF)^TDF,
\]

(3)

where \( D \) stands for differentiation with respect to \( x_0 \). For any smooth \( v, F \) represents a diffeomorphism, which ensures invertibility of \( DF \) and thus positive definiteness of \( C \). Furthermore, incompressibility of \( v \) implies \( \det C = 1 \). Consequently, eigenvalues \( \{\lambda_i\} \) and normalized eigenvectors \( \{\xi_i\} \) of \( C \) satisfy

\[
0 < \lambda_1 \leq \lambda_2 \equiv \frac{1}{\lambda_1}, \quad \xi_i \cdot \xi_j = \delta_{ij} \quad i, j = 1, 2.
\]

(4)

B. Coherent Lagrangian eddies

Haller and Beron-Vera (2013) seek elliptic LCS as material loops with small annular neighborhoods showing no leading-order variation in averaged material stretching (Fig. 2).

Solving this variational problem reveals that elliptic LCSs are uniformly stretching: any of their subsets are stretching after advection. Any of their subsets are uniformly stretching: any of their subsets are stretching after advection.

\[
0 \leq \dot{\lambda}_1 < \dot{\lambda}_2 \eqqquad \xi_i \cdot \xi_j = \delta_{ij} \quad i, j = 1, 2.
\]

(5)

where the prime stands for \( s \) differentiation. More geometrically, limit cycles of (5) are closed null geodesics of the metric tensor \( C - \lambda^2 I_d \), which is Lorentzian in the domain satisfying \( \lambda_1 < \lambda_2 < \lambda_2 \). This provides a relativistic interpretation of coherent Lagrangian eddies (for details, cf. Haller and Beron-Vera, 2013, 2014).

The limit cycles of (5) will either grow or shrink under changes in \( \lambda \), forming smooth annular regions of nonintersecting loops. The outermost member of such a band of coherent Lagrangian loops will be observed physically as the boundary of the coherent Lagrangian eddy. We refer to these maximal elliptic LCSs as coherent Lagrangian eddy boundaries.

Limit cycles of (5) tend to exist only for \( \lambda \approx 1 \). Material loops characterized by \( \lambda = 1 \) resist the universally observed material stretching in turbulence: they resume their initial arc length at time \( t \). This conservation of arc length, along with the conservation of the enclosed area in the incompressible case, creates extraordinary coherence for elliptic LCS.

C. Inertial effects near coherent Lagrangian eddies

The Maxey–Riley equation (Maxey and Riley, 1983) describes the motion of inertial (i.e., buoyant, finite-size) particles, which can deviate substantially from that of fluid (i.e., neutrally buoyant, infinitesimally small) particles (cf. Cartwright et al., 2010). Here, we consider a simplified version of the Maxey–Riley equation appropriate for inertial particle motion in a quasigeostrophic flow. We further derive a reduced form of this equation, which will allow us to assess behavior near a coherent Lagrangian eddy.

Specifically, ignoring added mass effects, the Basset history term, and so-called Faxen corrections, the Maxey–Riley equation for the motion of a small spherical particle in the flow produced by (1) is given by

\[
\ddot{x} + f \dot{x} = \frac{\rho}{\rho^p} \dot{\delta} \cdot v - \tau^{-1}(\dot{x} - v),
\]

(6)

where the constants

\[
\delta := \frac{\rho}{\rho^p}, \quad \tau := \frac{2a^2}{9D_\theta}.
\]

(7)

Here, \( \rho \) and \( \nu \) are the fluid’s density and kinematic viscosity, respectively, and \( \rho^p \) and \( a \) are the inertia particle’s density and radius, respectively. The left-hand-side of (6) is the inertial particle’s absolute acceleration. The first and second terms on the right-hand-side of (6) are the flow force and Stokes drag, respectively.

The simplified form of the Maxey–Riley equation (6) was priorly considered by Provenzale (1999) with the following differences. First, the fluid relative acceleration, \( \rho (\partial_\nu E + v \cdot \nabla v) \), was included. This term is one order of magnitude smaller in Ro than \( f \dot{v} \) and thus is conveniently neglected here. Second, a centrifugal force term was included too, but this is actually balanced by the
gravitational force on the horizontal plane. Third, a vertical buoyancy force term was considered, but this in the end played no role as the focus was on motion on a horizontal plane, as here.

We introduce the small nondimensional parameter

\[ \varepsilon := \tau \frac{V}{L} = \frac{2}{9d} \left( \frac{a}{L} \right)^2 \frac{\text{Re}}{St} \leq 1, \]

(8)

where \( \text{Re} \) and \( St \) are Reynolds and Stokes numbers, respectively. Consistent with the quasigeostrophic scaling assumptions leading to the fluid velocity field (1), we can set

\[ \varepsilon = O(\text{Ro}). \]

(9)

In Appendix A, we show that inertial particle motion characterized by (9), e.g., the motion of particles much smaller than the typical length scale of the flow, is controlled at leading order by

\[ \dot{x} = v^p = v + (\delta - 1) f v^\perp, \]

(10)

which is the reduced form of the Maxey–Riley equation we shall use. This reduced equation is valid up to an \( O(\varepsilon^2) \) error, after particles reach the vicinity of an attracting slow manifold exponentially fast.

Comparison of (2) and (10) reveals that the fluid velocity, \( v \), and inertial particle velocity, \( v^p \), differ by a dissipative \( O(\varepsilon) \) term. In the northern hemisphere \( (f > 0) \), this term acts to deflect the motion of positively buoyant \( (\delta > 1) \) finite-size or light particles to the left of the motion of fluid particles, while it acts to deflect the motion of negatively buoyant \( (\delta < 1) \) finite-size or heavy particles to the right; in the southern hemisphere \( (f < 0) \), it acts the opposite way (Fig. 3).

Inertial effects, therefore, should promote divergence away from, or convergence into, coherent Lagrangian eddies when otherwise fluid particles circulate around them.

Specifically, let \( \gamma \) be the boundary of a coherent Lagrangian eddy at time \( t \) and \( U_c \), the region \( \gamma \) encloses. Up to an \( O(\varepsilon^2) \) error, the flux across \( \gamma \) is given by

\[ \Phi_c = \int_{\gamma} (v - v^p) \cdot d\mathbf{x} = \int_{U_c} \nabla \cdot (v^p - v) \, d^2x = \tau (1 - \delta) \int_{U_c} \omega \, d^2x, \]

(11)

where the loop integral is taken anticlockwise and \( \omega := -\nabla \cdot v^p = (\alpha f^{-1} \nabla^2 \eta) \) is the fluid’s vorticity. Inspection of expression (11) leads to the following conclusions:

1. cyclonic \( (f_\omega > 0) \) coherent Lagrangian eddies attract \( (\Phi_c < 0) \) light \( (\delta > 1) \) particles and repel \( (\Phi_c > 0) \) heavy \( (\delta < 1) \) particles;

2. anticyclonic \( (f_\omega < 0) \) coherent Lagrangian eddies attract \( (\Phi_c < 0) \) heavy \( (\delta < 1) \) particles and repel \( (\Phi_c > 0) \) light \( (\delta > 1) \) particles.

Our results concerning heavy particles confirm the numerical observations of Provenzale (1999) and extend them to the behavior of light particles.

Our computations below are based on the reduced Maxey–Riley equation (10), which we refer to as the inertial equation. This follows the terminology of Haller and Sapsis (2008), who obtained the reduced form of a system similar to (6) in a nonrotating frame. Considering (10) is advantageous computationally and, as we will show, sufficiently accurate for the verification of our theoretical results.

D. Inertial Lagrangian coherent structures

While motion of inertial particles is not constrained by LCS, it is tied to analogous exceptional invariant curves referred to as inertial LCS (or iLCS; cf. Haller and Sapsis, 2008).

Of particular interest for our purposes here are hyperbolic iLCSs of attracting type. These can be obtained by applying recent LCS theory results (Haller and Beron-Vera, 2012; Farazmand and Haller, 2013; Farazmand et al., 2014; and Haller, 2015) on system (10). Specifically, iLCSs at time \( t_0 \), which attract nearby inertial particle trajectories over \([t_0, t]\), are invariant curves \( s \mapsto r(s) \) that satisfy

\[ r' = v^p_s \quad \text{or} \quad r' = v^p_g \]

(12)

and

\[ \sqrt{\lambda^2} > 1 \quad \text{or} \quad \sqrt{\xi^2} < 1 \]

(13)

for \( t < t_0 \) or \( t > t_0 \), respectively. Here, \( \{\lambda^p\} \) and \( \{\xi^p\} \) are eigenvalues and eigenvectors, respectively, of the Cauchy–Green tensor, \( C^p \), derived from system (10), which is an objective measure of deformation in that system. In forward time, segments of these invariant lines squeeze and

![FIG. 3. Velocity contributions to inertial particle's velocity: light (heavy) particle motion deflects to the left (right) of fluid particle motion in the northern hemisphere and vice versa in the southern hemisphere.](image)
stretched, respectively. As a result, they can be referred to as inertial squeeze lines and inertial stretch lines, respectively. In a similar manner as the $\lambda$-lines discussed above, these invariant lines admit a null geodesic interpretation. In this case, the relevant Lorentzian metric tensor is given by $C^p \Omega - \Omega C^p$, where $\Omega$ is a $90^\circ$-anticlockwise rotation matrix (Farazmand et al., 2014).

III. SIMULATIONS

Here, we present numerical results that confirm our theoretical predictions for the motion of inertial particles near coherent Lagrangian eddies in the ocean.

In each of our numerical tests, a coherent Lagrangian eddy was detected assuming that fluid trajectories are governed by (2) with the velocity field given in (1); the SSH field is constructed using satellite altimetry measurements (Fu et al., 2010). All eddies were detected from 90-day forward integration and found to have $\lambda = 1$. Successive positions of the boundaries of the eddies past the detection time were obtained from advection. Inertial particles were assumed to have $a = 0.25$ m, which is a realistic radius value for commonly employed spherical drifting buoys. Both light and heavy particles were considered, with $\delta = 1.1$ and 0.9, respectively. For typical oceanic mesoscale eddies, with diameter $L \sim 150$ km and tangential velocity at the boundary $V \sim 0.1$ m s$^{-1}$, these inertial particle parameter choices give $\varepsilon \sim 0.01$. This $\varepsilon$ value turned out to be small enough for particle motion obeying the Maxey–Riley equation (6) to exhibit behavior qualitatively similar to that satisfying the inertial equation (10) employed in our simulations.

We begin by discussing the results of tests involving light and heavy particles initially located on the same position on the boundary of a coherent Lagrangian eddy. The results are summarized in Fig. 4, which consider a cyclonic (left panel) and an anticyclonic (right panel) eddy, both indicated in light blue. The eddy in the left panel is identifiable with a cold-core Gulf Stream ring, while that in the right panel with an Agulhas ring. The arc length of the boundary of each eddy on the detection date is reassumed 90 days after (recall that the eddies have $\lambda = 1$). Coherence is nevertheless observed well beyond 90 days consistent with previous analyses of the satellite altimetry data set (Beron-Vera et al., 2013 and Haller and Beron-Vera, 2013). This is evident from the complete absence of filamentation. The light (green) and heavy (red) particles behave quite differently than the fluid particle (yellow) initially lying on the same position as the inertial particles on the boundary of each eddy. The fluid particle remains on all dates shown on the boundary of the Lagrangian eddy carrying the particle. Consistent with our predictions, the light (heavy) particle spirals into (out of) the cold-core Gulf Stream ring, while it spirals out of (into) the Agulhas ring.

We now provide more explicit support to our predictions by presenting the results from the computation of the pointwise flux of inertial particles across the boundary of a coherent Lagrangian eddy. Across a material loop $\gamma$, the pointwise flux of inertial particles is given by $\left( (v^p - v) \cdot n_c \right)$, where $n_c$ is the outer unit normal to $\gamma$. Taking $\gamma$ as the boundary of the eddy identified above as an Agulhas ring, the latter is plotted in Fig. 5 on 24 November 2006 as a function of the boundary parameter $s$, chosen to be an azimuthal angle. The pointwise fluxes of light (solid green) and heavy (solid red) particles are everywhere inward and outward, respectively, along the boundary of the anticyclonic eddy in question. Thus, our sign predictions for the total flux extend to the pointwise flux in this example.

We now turn to illustrate in Fig. 6 that the evolution of inertial particles is tied to attract iLCS. This is done for patches of light (green) and heavy (red) particles lying initially outside the coherent Lagrangian Agulhas ring discussed above (light blue). Shown attracting iLCSs (black) were computed as most stretching inertial stretch lines through each patch. This was done on the eddy detection time from a 90-day-forward integration. The evolution of each inertial stretch line was determined by advection. After experiencing substantial stretching, the light particle patch is repelled away from the eddy. By contrast, the heavy particle patch spirals into the eddy. As expected, the attracting iLCS forms the centerpieces of the patch in each case. For completeness, the evolution of a fluid patch (yellow) is also shown. In this case, too, the patch evolution is tied to its

FIG. 4. Simulated trajectories of light (green), heavy (red), and fluid (yellow) particles initially on the boundaries of two mesoscale coherent Lagrangian eddies (light blue) extracted from altimetry-derived velocity. Advection for the fluid particles is supplied by the altimetry-derived velocity, and heavy and light particle motion is controlled by the inertial equation (10). The eddies are identifiable with a cyclonic cold-core Gulf Stream ring (left panel) and an anticyclonic Agulhas ring (right panel).
centerpiece attracting LCS, computed also as the most stretching stretchline through patch. Consistent with the material nature of the boundary of the coherent eddy, the fluid patch spirals around the eddy without penetrating it.

Finally, we show that for the parameters chosen, the qualitative behavior near a coherent Lagrangian eddy described by the Maxey–Riley equation is captured by the inertial equation. This is illustrated in Fig. 5, where the dashed lines correspond to pointwise flux calculations based on the inertial and Maxey–Riley equations, respectively. Once again, while details of inertial and Maxey–Riley-equation-based trajectories are different and convergence can only be expected when particles are heavy, our predictions are seen to hold well: the heavy particle is attracted by the anticyclonic coherent Lagrangian eddy in question, whereas the light particle is repelled away from it.

IV. OBSERVATIONS

In this section, we discuss four sets of ocean observations that can be explained using our predictions for the motion of inertial particles near mesoscale coherent Lagrangian eddies.

The first set of observations, discussed in the Introduction, concern two RAFOS floats in the southeastern North Pacific. The floats took divergent trajectories despite their initial proximity within an anticyclonic mesoscale eddy. Revealed as a region of closed altimetric SSH streamlines, this eddy may be the Eulerian footprint on the SSH field of a coherent Lagrangian California Undercurrent eddy or “cuddy.” In effect,
as we anticipated in the Introduction and now explicitly show
in the top panel of Fig. 8, the altimetry-derived velocity field
sustains a coherent Lagrangian eddy in nearly the same posi-
tion as the SSH eddy. The eddy, obtained from a 90-day-for-
ward integration with $k = 1$, is depicted (in light blue) on the
detection date and two subsequent dates. The trajectory of each
of the two floats is indicated by a curve, with their initial posi-
tion highlighted by a dot. The divergent behavior of the float
trajectories can be explained by inertial effects as follows.

The bottom-left panel of Fig. 8 reveals that the float
indicated in green experiences a net ascending motion from
24 July 2003, the beginning of the record, through about 10
October 2003, roughly when the float escapes the coherent
Lagrangian eddy, detected from altimetry on 21 August
2003. By contrast, the bottom-right panel of Fig. 8 shows
that the float indicated in red oscillates about a constant
depth over this period but experiences a net descending
motion from 10 October 2003 till the end of the observatio-
nal record, 18 March 2004. Positive overall buoyancy can
thus be inferred for the green float from the beginning of the
observational record until about 10 October 2003. By con-
trast, negative overall buoyancy, preceded by a short period
of neutral overall buoyancy, can be inferred for the red float
over the entire observational record. Appendix D provides
quantitative confirmation for these qualitative inferences
and further shows that these are not sensitive to the time
windows considered.

The sign of the overall buoyancy of each float can be
used to describe its behavior qualitatively through the inertial
equation (10). The green float remains within the anticy-
clonic coherent Lagrangian eddy from 21 August to around
10 October 2003, nearly when it leaves the eddy and does
not come back during the total observational record (about 6
months). This is qualitatively consistent with the behavior of
a light particle. Beyond 10 October 2003, the buoyancy sign
for this float is not relevant, given that it is already outside
the eddy. In contrast, the red float remains inside the
Lagrangian eddy over the whole observational record.
This is qualitatively consistent with the behavior of a heavy
particle. We note that there are shorter, intermittent periods
of positive buoyancy for the red float within this observatio-
nal record. The float, however, does not have to leave the
eddy during these periods. Rather, these intervals signal a
nonuniformity in the stability for the underlying eddy core,
which nevertheless prevails as a net attractor over the obser-
vational period. Appendix D includes results from simula-
tions using the inertial equation, which provide support to
these qualitative inferences.

The second set of observations concerns trajectories of
satellite-tracked surface drifters deployed in the Gulf of

FIG. 7. Simulated trajectories of light (green) and heavy (red) particles ini-
tially lying on the same location on the boundary of the coherent Lagrangian
Agulhas ring of the previous figures (light blue) based on inertial (solid) and
Maxey–Riley (dashed) equations.

FIG. 8. (Top panel) Trajectories of the two RAFOS floats in Fig. 1 (green and
red) and snapshots of an anticyclonic coherent Lagrangian eddy detected
from altimetry (light blue). The dots indicate the positions of the floats on
the dates that the eddy is shown. (Bottom-left panel) Depth of the green
float as a function of time. (Bottom-right panel) As in the bottom-left
panel, but for the red float.
Mexico ahead of hurricane Rita in September 2005. The drifters were deployed inside a Loop Current ring detected from its Eulerian footprints in the altimetric SSH field. A 30-day-forward integration of the altimetry-derived velocity field reveals that the anticyclonic SSH eddy contains a $\lambda = 1$ coherent Lagrangian eddy with a radius of roughly 100 km, about 25-km smaller than that of the approximately circular area occupied by the SSH bulge. Figure 9 shows the coherent Lagrangian eddy (light blue) on the detection date and two subsequent positions obtained from advection. The trajectories of the drifters (a total of nine) are depicted in green (positions on the date shown are indicated by dots). Hurricane Rita made landfall about one week prior to the detection date, so neither the altimetry signal nor the motion of the drifters is affected by the high winds associated with this tropical cyclone system. Three drifters lie inside the eddy on the detection date, while the remaining six are located outside of the eddy, but close by its boundary. Overall, the drifters undergo growing looping trajectories. More than one month after the detection date, all nine drifters are found away from the center of the eddy, with three lying on its border and six lying well away from it. Noting that the drifters maintain afloat on the ocean surface, this behavior can be expected, given that anticyclonic coherent Lagrangian eddies repel away light particles according to our results.

The third set of observations involves the trajectory of a surface drifters tracked by the Argos satellite system, which was deployed inside an Agulhas ring, named Astrid, as part of the Mixing of Agulhas Rings Experiment (MARE) \cite{van Aken et al., 2003}. (Two additional drifters were deployed during MARE whose trajectories have not been possible to access. However, all three drifters behave similarly as it can be seen in Fig. 6 of \cite{van Aken et al., 2003}.) Detected from its Eulerian footprints in the altimetric SSH field, ring Astrid was subjected to a detailed survey. Hydrographic casts across ring Astrid indicated the presence of a warm and saline core. Acoustic Doppler current profiling revealed that Astrid had, in addition to the baroclinic flow around its core, a significant barotropic component. A 30-day forward integration of the altimetry-derived velocity field reveals a coherent Lagrangian eddy with $\lambda = 1$. This eddy has a mean radius of roughly 100 km, about half that of the approximately circular region spanned by the SSH bulge. Figure 10 shows selected snapshots of the coherent Lagrangian eddy (light blue) on three dates starting from the detection date. The trajectory of the Argos-tracked surface drifter is indicated in green (dots indicate the positions of the drifter on the corresponding dates). The drifter is seen to develop counterclockwise looping trajectories. This grows in radius and quite quickly the drifters abandon the vicinity of the eddy. The coherent Lagrangian eddy is rather short lived, thereby not revealing the presence of a well-developed Agulhas ring (possibly consistent with the lack of a well-defined core in the \textit{in-situ} velocity measurements). However, over the lifespan of the eddy, the drifters’ behavior is consistent with our predictions for a light particle. Therefore, our results offer an explanation for its motion.

Finally, the fourth set of observations involves distribution of floating \textit{Sargassum} on the sea surface in the western North Atlantic inferred from the Medium Resolution Imaging Spectrometer (MERIS) aboard \textit{Envisat} (Fig. 11, top panel). \textit{Sargassum} corresponds to Maximum Chlorophyll Index (MCI) values exceeding $-0.25 \text{ mW m}^{-2} \text{ sr}^{-1} \text{ nm}^{-1}$. Detected on 4 October 2006, the feature of interest takes a spiraled shape and lies inside a coherent Lagrangian cold-core Gulf Stream ring as revealed from altimetry. In the bottom-left panel of Fig. 11, the material boundary of this cyclonic coherent ring is shown overlaid on the \textit{Sargassum} feature in question. This was obtained in Section \textbf{III} from advection of a coherent Lagrangian eddy boundary.
computed on 5 August 2006 from a 90-day-forward integration of the altimetry-derived velocity field (cf. Fig. 4, top-right panels). The accumulation of the floating Sargassum inside the Gulf Stream ring is consistent with the behavior of inertial particles near cyclonic coherent Lagrangian eddies, which attract light particles according to our results. The spiraled shape of the Sargassum distribution inside the ring is consistent too with the spiraled shape acquired by altimetry-based attracting light iLCS (parameters are as in Section III). Selected iLCSs are shown overlaid on the Sargassum distribution in the bottom-right panel of Fig. 11. These were obtained as backward-time light inertial squeezelines initialized along the boundary of the Gulf Stream ring on the date shown. The direction of the spiraling inertial particle motion along these iLCS is inward, as direct integration of the inertial particle equation reveals.

V. SUMMARY AND DISCUSSION

In this paper, we have provided an explanation for the observed tendency of drifting buoys and floating matter on the surface of the ocean to produce dissipative-looking patterns. This resolves an apparent paradox with the conservative-looking distributions that tracers passively advected by a rotating two-dimensional incompressible flow display. Our explanation takes into account inertial effects, i.e., those produced by the buoyancy and size finiteness of an object immersed in such a flow. These are described by a simplified Maxey–Riley equation consistent with a flow produced by a quasigeostrophic velocity where the pressure field is entirely due to differences in sea surface height. Because the latter are readily available from satellite altimetry measurements, our approach enables feature matching and analysis of specific observations. Furthermore, our approach is self-consistent within the realm of two-dimensional incompressible flows.

We have found that anticyclonic coherent Lagrangian eddies attract (repel) heavy (light) particles, while cyclonic ones behave the opposite way. We verified these results numerically using mesoscale SSH fields constructed from satellite altimetry measurements in various places of the ocean. Our findings also explained dissipative-type behavior shown by four sets of observations: divergent motion of subsurface floats initially inside a California Undercurrent eddy or “cuddy,” dispersion of surface drifters away from a Loop Current ring, ejection of surface drifters out of a well-studied Agulhas ring, and accumulation of sargassum inside of a cold-core Gulf Stream ring.

Beyond the reach of the Maxey–Riley description is motion of arbitrarily shaped objects; no known theory accounts for their effects. At the Maxey–Riley level, there are terms and aspects that we have ignored, which may contribute to narrow the gap between theory and observed motion. One such term is the memory term, but this only tends to slow down the inertial particle motion without changing its qualitative dynamics fundamentally (Daitlech and Tél, 2011). Another neglected aspect is the dependence of fluid density on spatial position and time. Time varying density effects were investigated previously in idealized settings and found to be of importance (Tanga and Provenzale, 1994). The observational possibility to account for these effects is provided by satellite sensing of sea surface temperature and salinity. An additional aspect is the effect of submesoscale perturbations on the mesoscale motions of interest to us here. These may be of fully ageostrophic and possibly three-dimensional nature (McWilliams, 2008) or still be balanced to leading order, and thus essentially two-dimensional and incompressible (Klein and Lapeyre, 2009).

The latter is particularly interesting as it opens the way to a potentially more accurate observation-based velocity representation when high-resolution wide-swath altimetry becomes operational (Fu and Ferrari, 2009). The observational improvement over altimetry-derived velocities may then be expected from the addition of an Ekman drift component estimated from satellite scatterometer wind measurements. This typically small correction is regularly included, but in such a way as to match observed drifting buoy velocities (Lagerloef et al., 1999), which is not consistent with our inertial particle approach.

We finally note that a larger sample of drifting buoys and floating matter than that considered here is required to further validate our predictions, possibly improved by the consideration of fluid density variations and Ekman drift effects.

FIG. 11. (Top panel) Maximum Chlorophyll Index (MCI) in the western North Atlantic inferred from the Medium Resolution Imaging Spectrometer (MERIS) aboard Envisat on 4 October 2006. Floating Sargassum corresponds to MCI values in excess of ~0.25 mW m^{-2} sr^{-1} nm^{-1}. (Bottom-left panel) Boundary of a coherent Lagrangian cold-core Gulf Stream ring detected from altimetry (black) overlaid on the Sargassum distribution. (Bottom-right panel) Altimetry-based attracting light iLCS (black) overlaid on the Sargassum distribution.
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APPENDIX A: REDUCED MAXEY–RILEY EQUATION

The second-order ODE (6) is equivalent to the following first-order ODE set:

\[ \dot{x} = v^p, \quad \dot{v}^p = f(\delta v - v^p) + \tau^{-1}(v - v^p). \]  

(A1)

Taking \( L \) and \( L/V \) as length and time scales, respectively, and making \( \tau f \equiv 1 \) so that \( \epsilon = \tau V / L = \text{Ro} \ll 1 \), the nondimensional form of (A1) reads

\[ \dot{x} = \epsilon v^p, \quad \epsilon^2 \dot{v}^p = \text{sign} f\epsilon(\delta v - v^p) + v - v^p. \]  

(A2)

Inspection of (A2) reveals that \( x \) is a slow variable that changes at \( O(1) \) speed, while \( v^p \) is a fast variable varying at \( O(\epsilon^{-2}) \) speed. Consequently, (A2) represents a singular perturbation problem. To regularize it, we displace and rescale time as \( \epsilon^{-2}(t - \lambda_0) \). Denoting with a circle differentiation with respect to this fast time variable, (A2) transforms into

\[ \dot{x} = \epsilon^2 v^p, \quad \epsilon^2 \dot{\epsilon} = \text{sign} f\epsilon(\delta v - v^p) + v - v^p, \quad \dot{\lambda} = \epsilon^2. \]  

(A3)

The \( \epsilon = 0 \) limit of system (A3) has a manifold of fixed points. This manifold is normally attracting and hence survives for small \( \epsilon > 0 \) in the form

\[ v^p = v + \epsilon v_1 + O(\epsilon^2). \]  

(A4)

Plugging this asymptotic series expansion into the right-hand-side equation of system (A3) and equating \( O(\epsilon) \) terms, it follows that

\[ v_1 = \text{sign} f(\delta - 1)v^p. \]  

(A5)

Inserting this expression in the left-hand-side equation of system (A3), the inertial equation (10) follows once dimensional variables are recovered.

Particle dynamics governed by the inertial equation (10) evolve, over the finite-time interval of interest, on a two-dimensional manifold, \( \mathcal{M}_1 \), in the phase space with coordinates \((x, v^p, t)\). This manifold is often referred to as slow because (10) restricted to it is a slowly varying system of the form \( \dot{x} = \epsilon^2 v^p|_{\mathcal{M}_1} = \epsilon^2 v + \epsilon^2 v_1 + O(\epsilon^4) \). As shown in Haller and Sapsis (2008), this slow manifold attracts all inertial particle motions exponentially.

APPENDIX B: DATA

The altimetric SSH data employed in this paper consist of background and perturbation components. The background SSH component is steady, given by a mean dynamic topography constructed from satellite altimetry data, in-situ measurements, and a geoid model (Rio and Hernandez, 2004). The perturbation SSH component is transient, given by altimetric SSH anomaly measurements provided weekly on a 0.25°-resolution longitude–latitude grid. This perturbation component is referenced to a 20-year (1993–2012) mean, obtained from the combined processing of data collected by altimeters on the constellation of available satellites (Le Traon et al., 1998). Mean dynamic topography and altimetry data are distributed by AVISO at http://www.aviso.oceanobs.com.

The RAFOS float trajectory data belong to the extensive data set constructed from float deployments in the California Undercurrent over the period 1992–2010 (Collins et al., 2013). Acoustically tracked, these RAFOS floats are quasi-isobaric, with their density varying with ambient temperature changes as a result of differing thermal expansions of the glass hull and aluminum end cap of the floats (Rossby et al., 1986). As opposed to seawater parcels, the floats sink when they warm and rise when they cool (Swift and Riser, 1994). The specific floats considered in this paper are shallow (300 db) floats, numbers 105 and 106, obtained from http://www.oc.nps.edu/npsRAFOS.

The surface drifters in the Loop Current ring were deployed from air by the 53rd Hurricane Hunter Squadron ahead of hurricane Rita. Equipped to monitor surface conditions, these drifters were of Minimet (drogue at 15 m) and ADOS (with a 100-m-long thermistor chain hanging below) types. The trajectories of these drifters are available from the NOAA Global Drifter Program at http://www.aoml.noaa.gov/phod/dac.

Three surface drifters were deployed in ring Astrid during the MARE-1 cruise. These were standard spherical WOCE/TOGA drifters, fitted with an 8-m-long holey sock drogue at 15 m, with their positions tracked using the Argos satellite system (Sybrandt and Niiler, 1991). The trajectories of these drifters are not available from any database. The trajectory of the drifter considered in this paper was digitalized from Fig. 4 of van Aken et al. (2003) and spline fitted. We have not been able to reliably digitalize from this figure the other two trajectories, which exhibit a qualitatively similar behavior.

Finally, the MERIS image shown in Fig. 11 is of L1b MCI, a spectrometer parameter traditionally used to detect and track Sargassum (Gower and King, 2008). MERIS imagery is available from ESA Earth Online at https://earth.esa.int/web/guest/data-access.

APPENDIX C: COMPUTATIONAL DETAILS

The flow maps associated with (2) and (10) were obtained from integration for initial positions on a regular 0.5-km-width grid covering the domain of interest. This was done using a stepsize-adapting fourth-order Runge–Kutta method with interpolations obtained using a cubic scheme. The derivative of the flow maps was computed using finite differences on an auxiliary 0.1-km-width grid of four points neighboring each point in the above grid. Integrations of (5) and (12) were carried out using the same method while enforcing a unique orientation for the corresponding vector fields at each integration step (recall that these are
constructed from eigenvector fields, which are not globally orientated). Detailed algorithmic steps for the extraction of coherent Lagrangian eddies are outlined in Haller and Beron-Vera (2013). The numerical computation of attracting LCS involves the same algorithmic steps as those for attracting LCS; these are outlined in Haller and Beron-Vera (2012) and Farazmand and Haller (2013).

APPENDIX D: EXTENDED ANALYSIS FOR THE RAFOS FLOAT DATA SET

As noted in Appendix B, the RAFOS floats considered here are designed to travel on isobaric surfaces, but these differ from constant depth surfaces due to ambient density changes, which induce small buoyancy variations for the floats. As a result, knowledge of the volume of a float is not sufficient to determine its buoyancy. To provide a rough quantitative confirmation for the sign of the float buoyancy, we assume a simplified model, whereby the vertical float motion decouples from (10) and is controlled by gravity and buoyancy, subject to Stokes drag. Under this approximation, the float’s vertical position, $z$, satisfies the equation of motion $\ddot{z} = (\delta - 1)g - \tau^{-1}z$, where $g$ is the gravity. The general solution to this equation is $z(t) = z_0 + (\delta - 1)gt$ $+ a_0 \exp(t-t_0) + a$ $+ \tau^{-1}2$, for some constants $z_0$ and $a_0$. Nonlinear regressions of this general solution to the observed floats’ depth histories are attained with $\delta - 1 \approx (0.48 \pm 0.13) \times 10^{-10}$ and $-(0.12 \pm 0.11) \times 10^{-10}$ for the green and red floats, respectively. These values were found by taking $a = 0.2$ m, the radius of a sphere whose volume matches that of a typical RAPOS float, which gives $\tau \approx 0.1$ d$^{-1}$. Uncertainties noted are of one standard deviation, resulting from considering regressions over time windows as short as 5 weeks (about twice the eddy turnover timescale) sliding over relevant observational time intervals. For the green float, the relevant observational time interval goes from the beginning of the record until the time the float escapes the eddy (11 weeks), while for the red float it is entire record (34 weeks). The black curves in the bottom panels of Fig. 12 correspond to regressions over the entire relevant time intervals. These results provide further confirmation to the visually inferred buoyancy signs, which we inferred only from the net vertical displacements of the floats between the endpoints of these intervals.

To match the actual shape of observed float trajectories, we would need to include further terms (such as nonlinear drag, added mass, horizontal–vertical coupling, etc.). These terms, however, would all bring in unknown parameters, thereby increasing the uncertainty in the inferred buoyancy. Instead, we proceed by solving (10) using a priori specified $\delta$ values, with the signs inferred above, but with magnitudes enhanced to a level that produces motions similar to the observations. Specifically, in the top panels of Fig. 12, we show simulated inertial particle trajectories of (10) with $\delta = 1.001$ (left) and 0.999 (right) starting from several initial positions inside the coherent Lagrangian eddy on 21 August 2003 rather than just the floats’ initial positions. This replaces a single trajectory model for the observed behavior with an ensemble of model trajectories. For both light and heavy particles, we have chosen (as above) $\tau = 0.1$ d$^{-1}$. This is roughly the inverse of the local Coriolis parameter, which makes $\epsilon$ small enough for (10) to hold asymptotically, as shown in Appendix A. The mean inertial particle trajectory inferred from the ensembles is indicated in each top panel of Fig. 12 by a thick curve. This ensemble averaged behavior is consistent with the observed float behavior, thereby justifying our buoyancy choices for modelling purposes. These choices might be improved by further optimization.

We close by noting that dealing with noisy observations is challenging and that while the behavior of the observed floats is consistent with the inertial equation derived here, more such observations with float pairs would be desirable to more thoroughly test the inertial equation predictions.


